

Image credit: Abel & Kaehler

NEUTRINO MASS FROM COSMOLOGY

THE HEBREW UNIVERSITY OF JERUSALEM

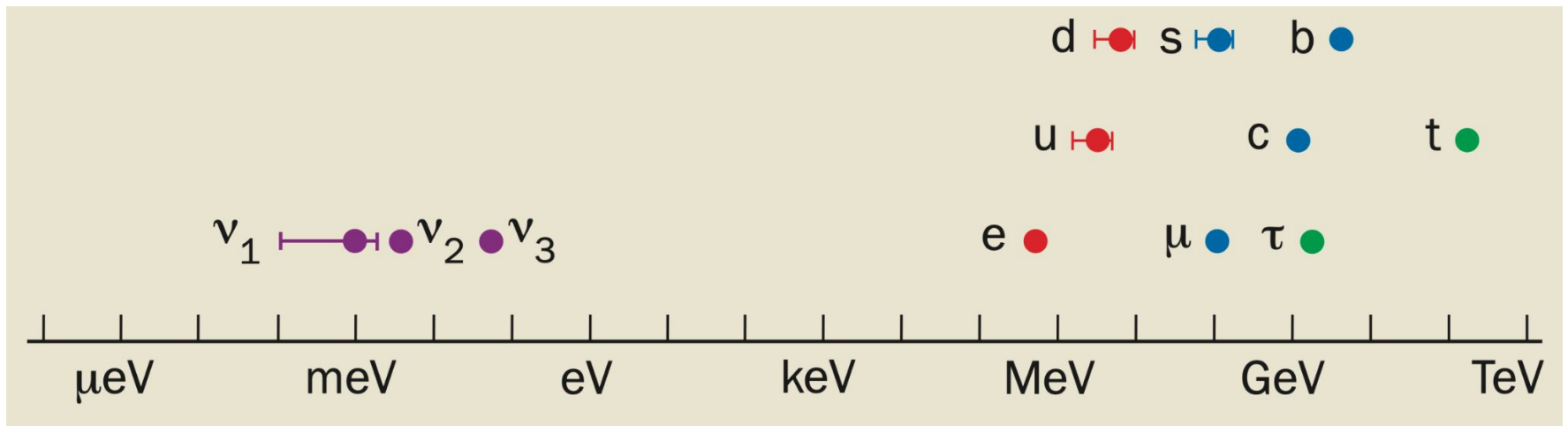
Unsolved Problems in Astrophysics and Cosmology, December 5, 2022

Jia Liu

KAVLI
IPMU

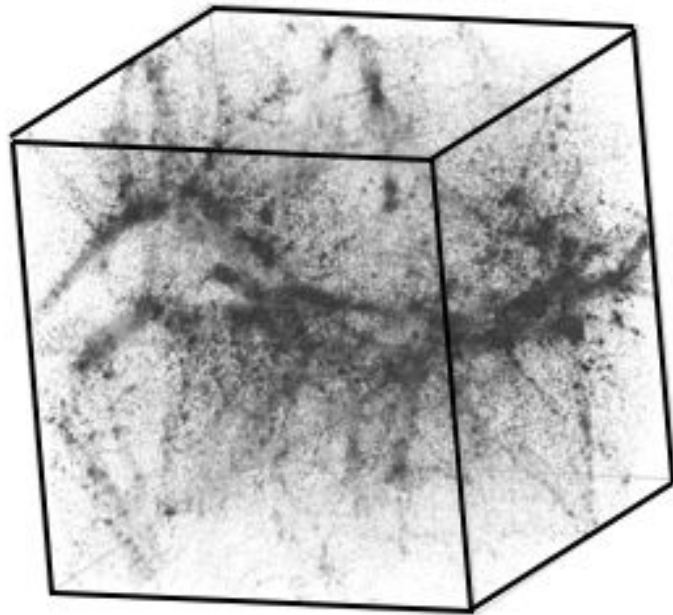
東京大学 国際高等研究所 カブリ数物連携宇宙研究機構
KAVLI INSTITUTE FOR THE PHYSICS AND MATHEMATICS OF THE UNIVERSE

Fermion masses

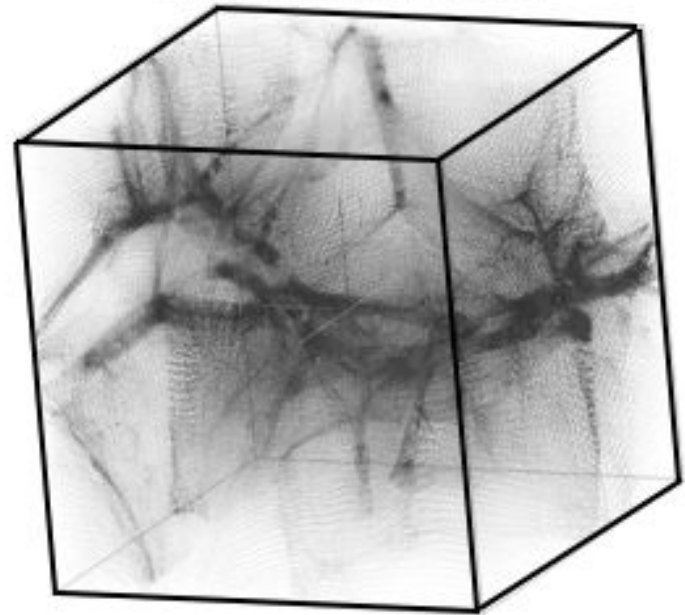


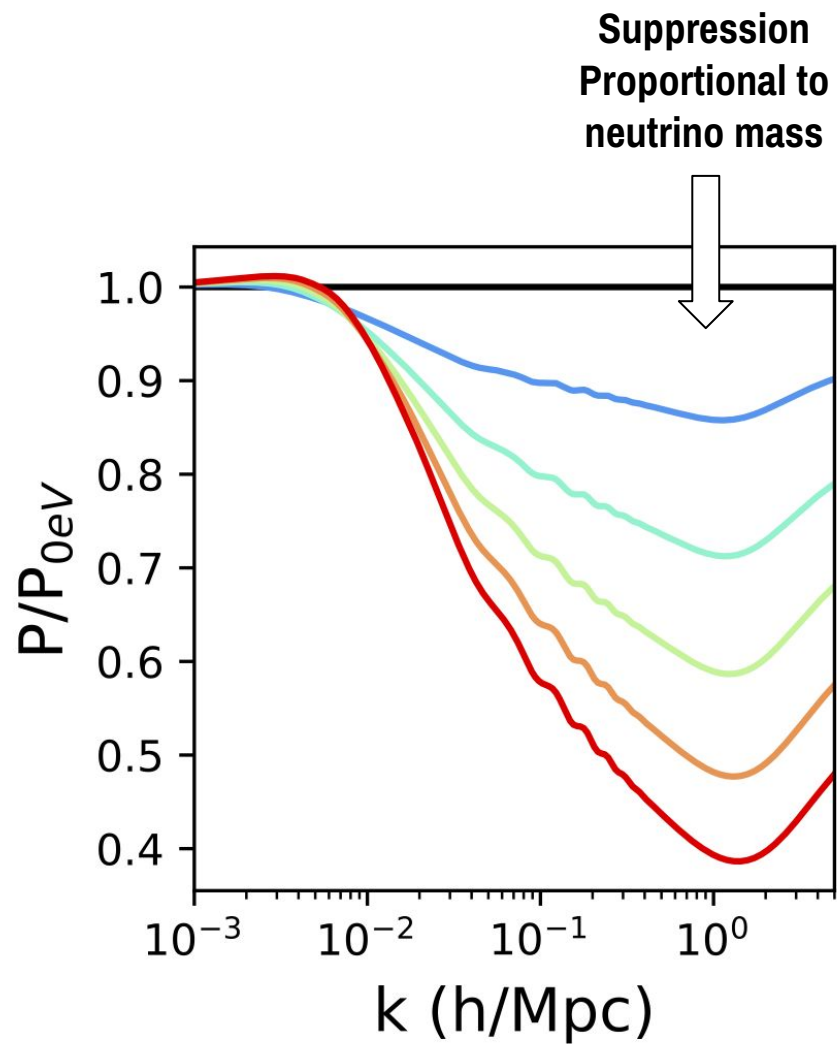
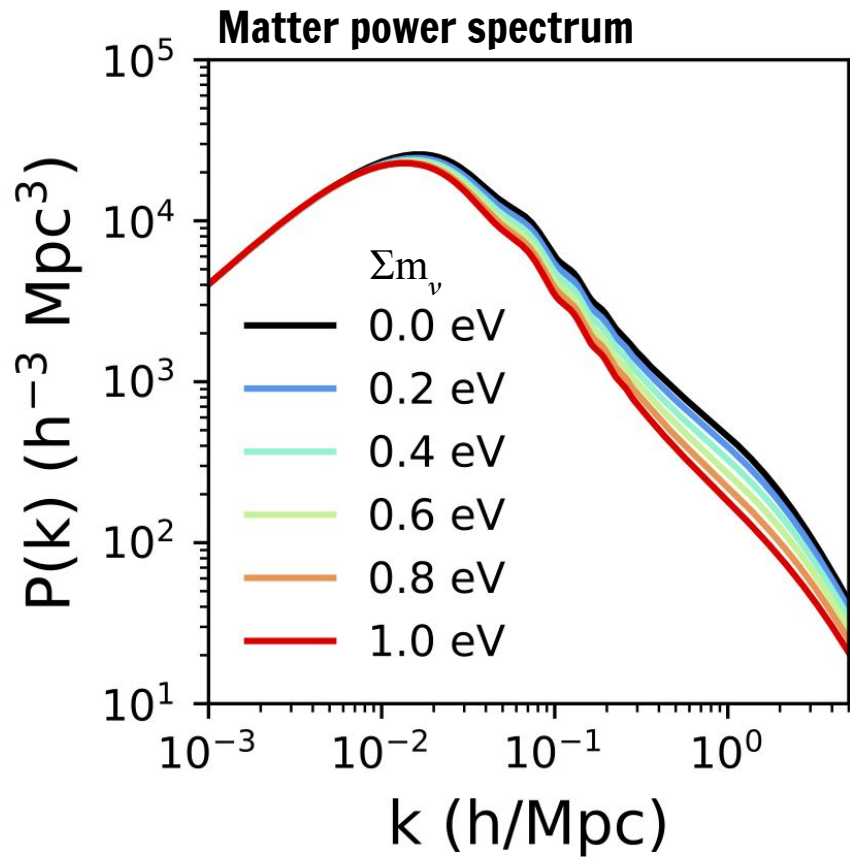
Hitoshi Murayama

Massless Neutrino



Massive Neutrinos





PARTICLE EXPERIMENT

COSMOLOGY

**Present
Upper Limit**

0.8 eV

KATRIN (2022, 90% CL)

0.13 eV

Planck + BAO
(2018, 95% CL)

**Future
Sensitivity**

~ 0.2 eV

KATRIN 2023

~ 0.02 eV

LSST, Euclid, PFS, DESI,
SO, CMB-S4, LiteBIRD

2018: 2 eV
(Troitsk)

2018: 0.23 eV
(Planck15+BAO)

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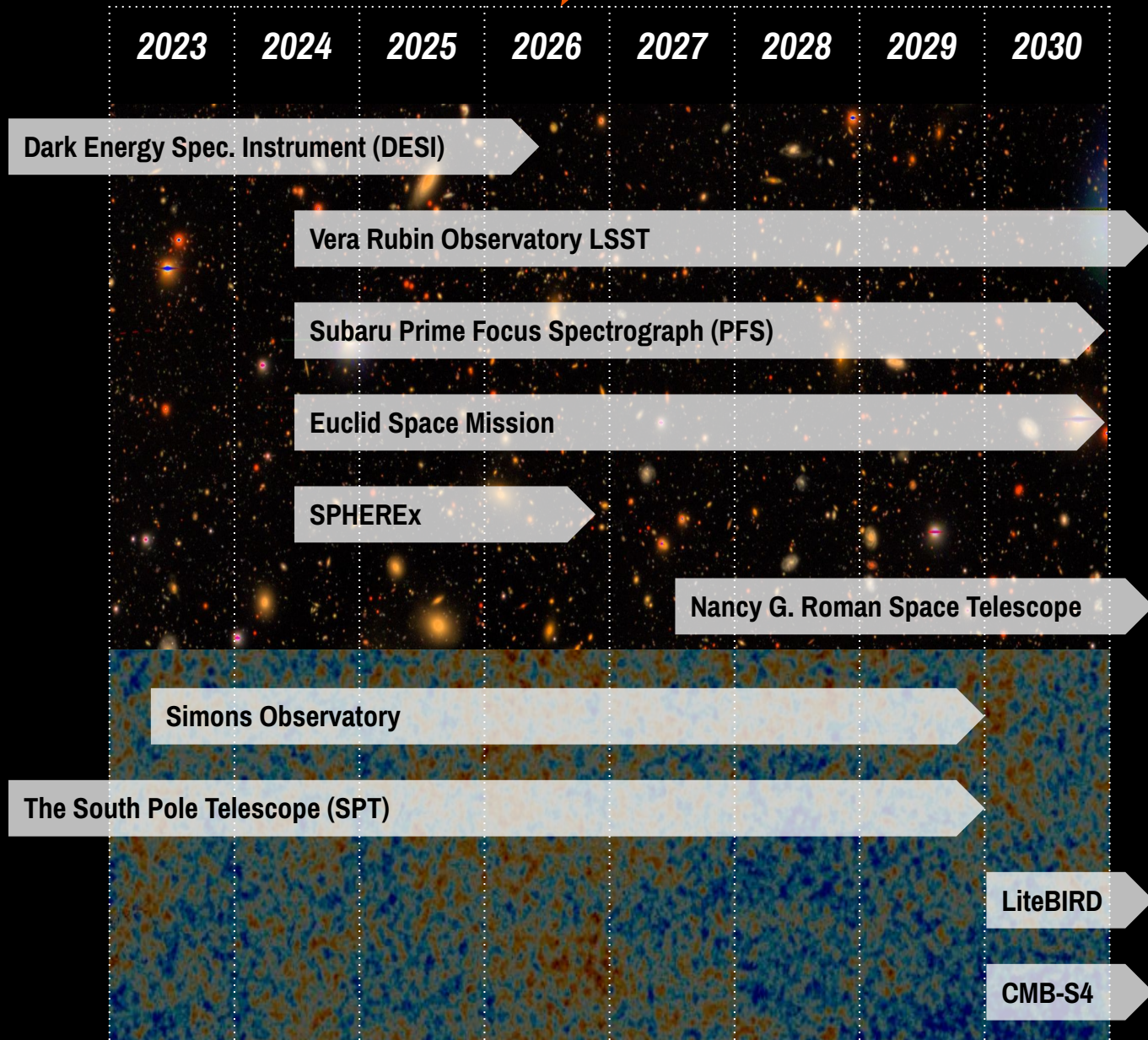
LSST, Euclid, PFS, DESI,
SO, CMB-S4, LiteBIRD

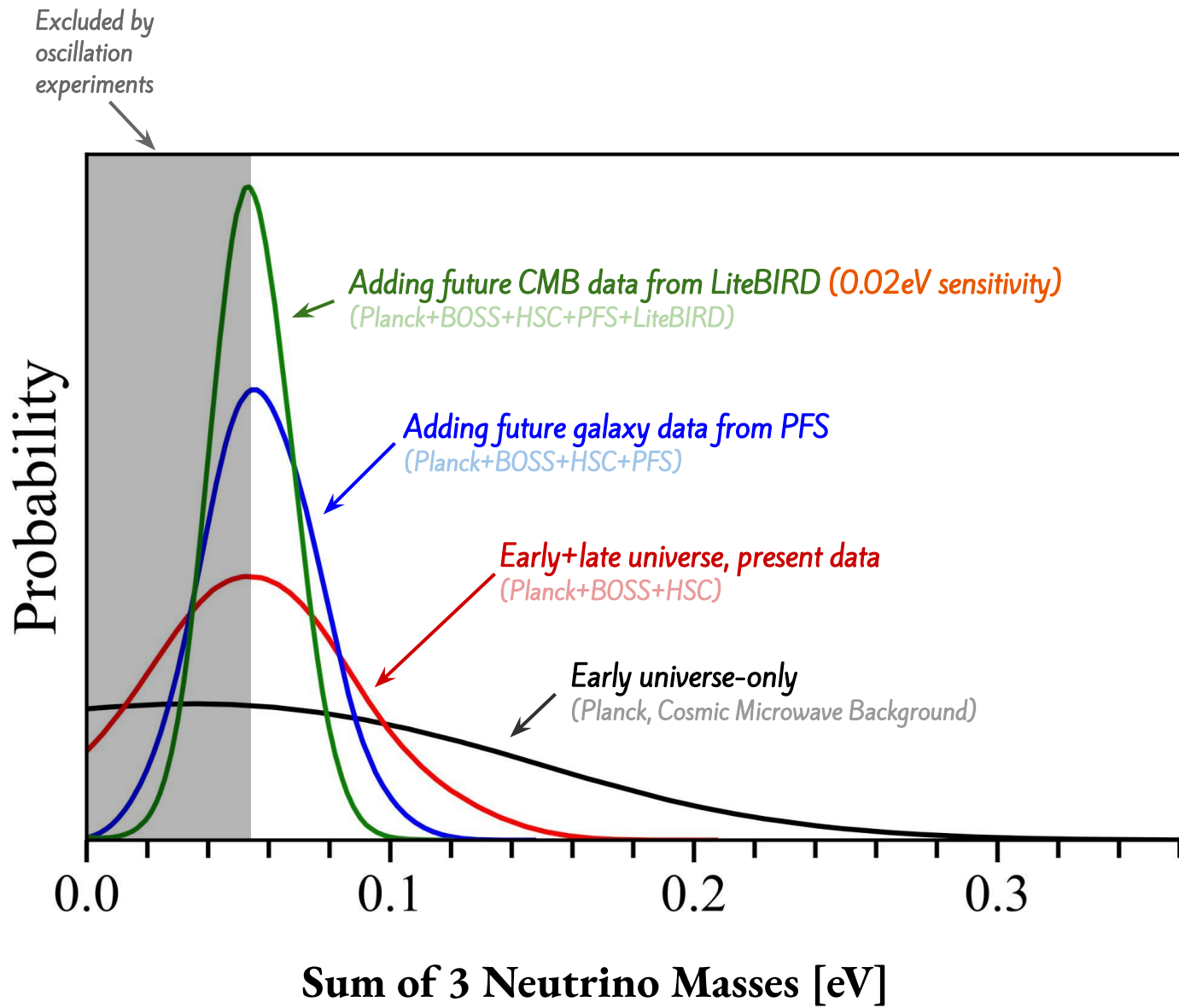
Normal
 $\Sigma m_\nu > 0.06 \text{ eV}$

Inverted
 $\Sigma m_\nu > 0.1 \text{ eV}$



Next "Unsolved Problems"





Normal
 $\Sigma m_\nu > 0.06 \text{ eV}$



Inverted
 $\Sigma m_\nu > 0.1 \text{ eV}$

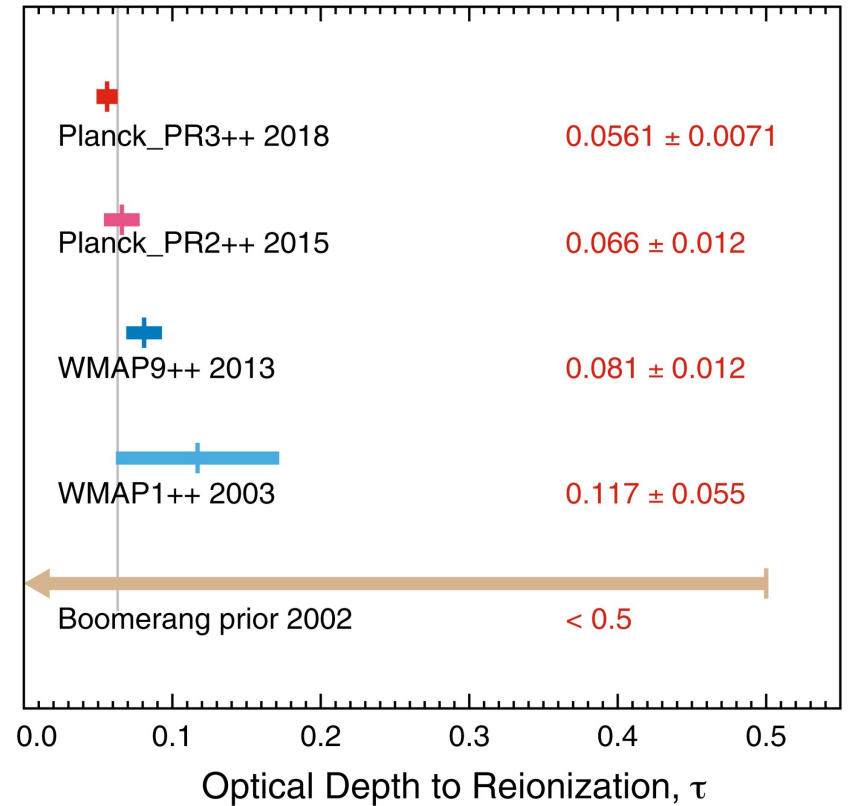
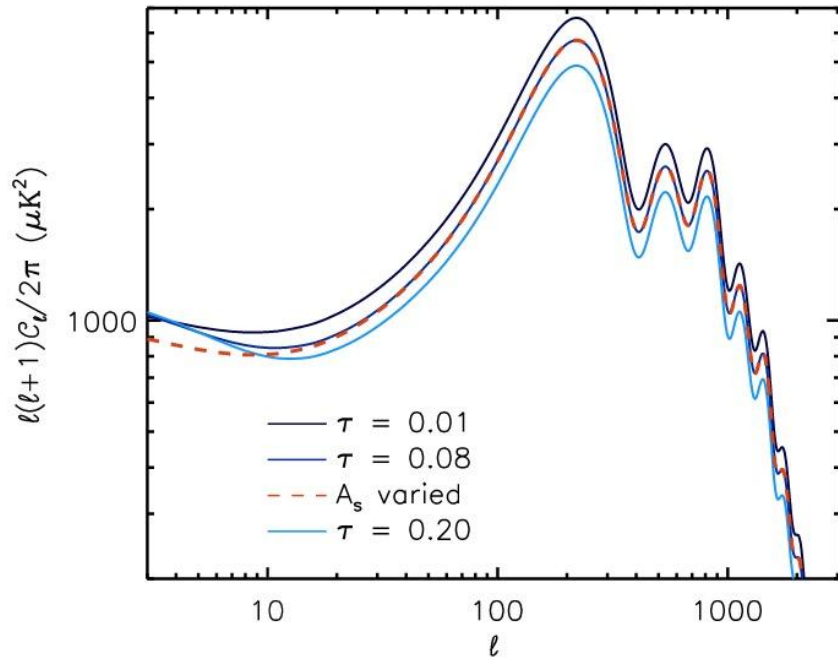


**0.02 eV sensitivity
= Discovery**

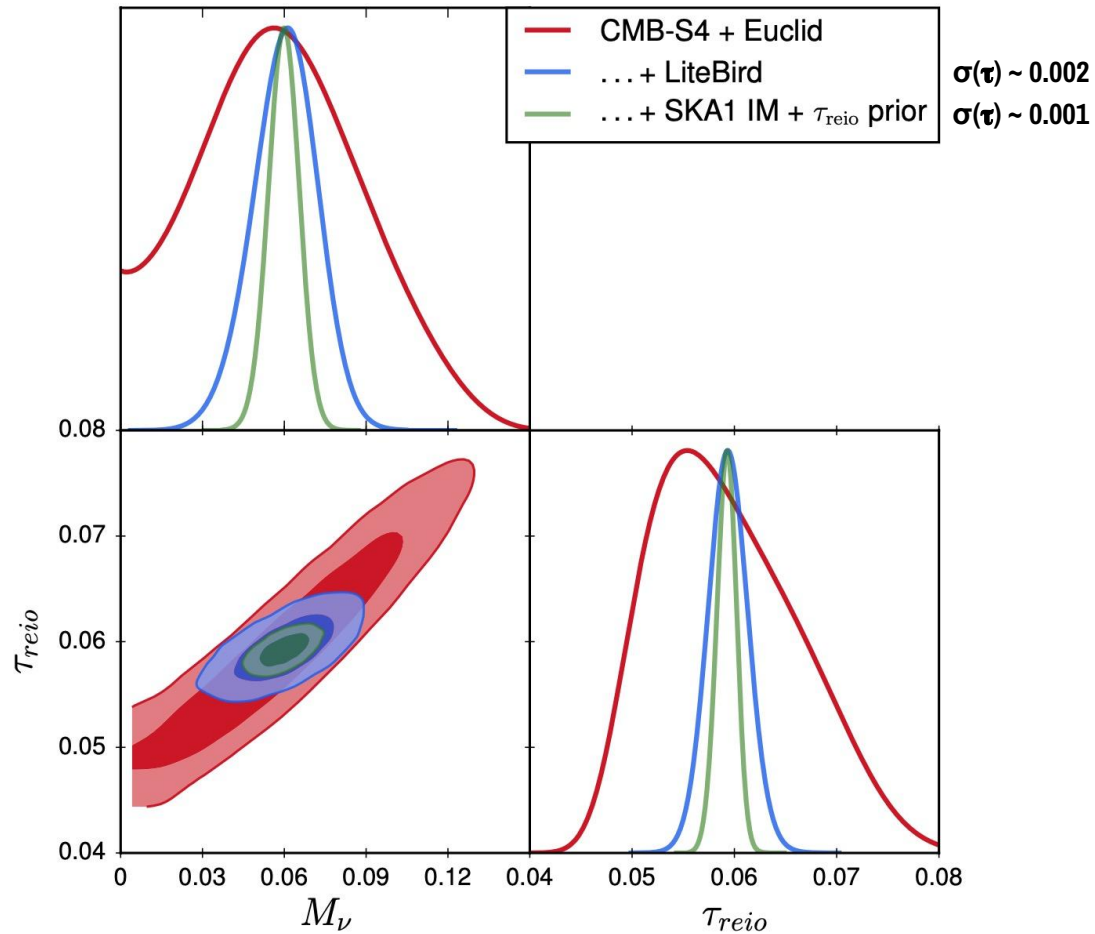
	Neutrino Mass	Mass Hierarchy
$\Sigma m_\nu > 0.1 \text{ eV}$	Discovery	No
$\Sigma m_\nu < 0.1 \text{ eV}$	Maybe	Discovery

Challenge 1: Optical Depth to Reionization τ_{reio}

Currently limited by Planck's τ_{reio}



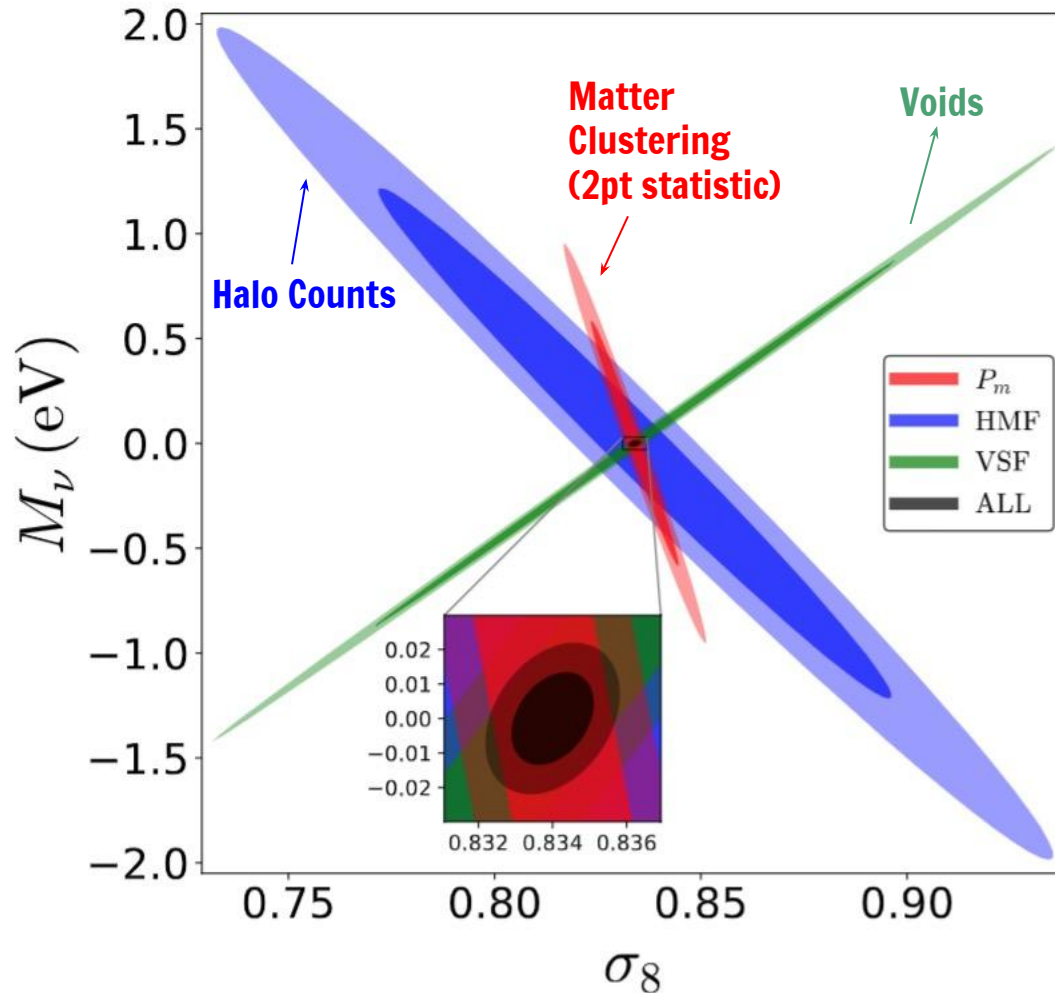
Better τ_{reio} from LiteBIRD (in 10 years)



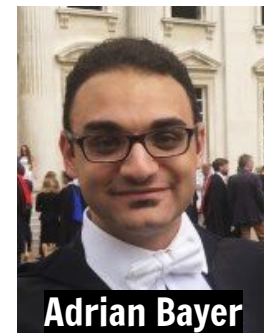
Brinckmann+2019

* assumes minimal 7 parameter Λ CDM+ M_ν model

Before LiteBIRD: Potential in Beyond-2pt Statistics



Requires significant improvement in theoretical modeling of nonlinear structure growth with numerical simulations.

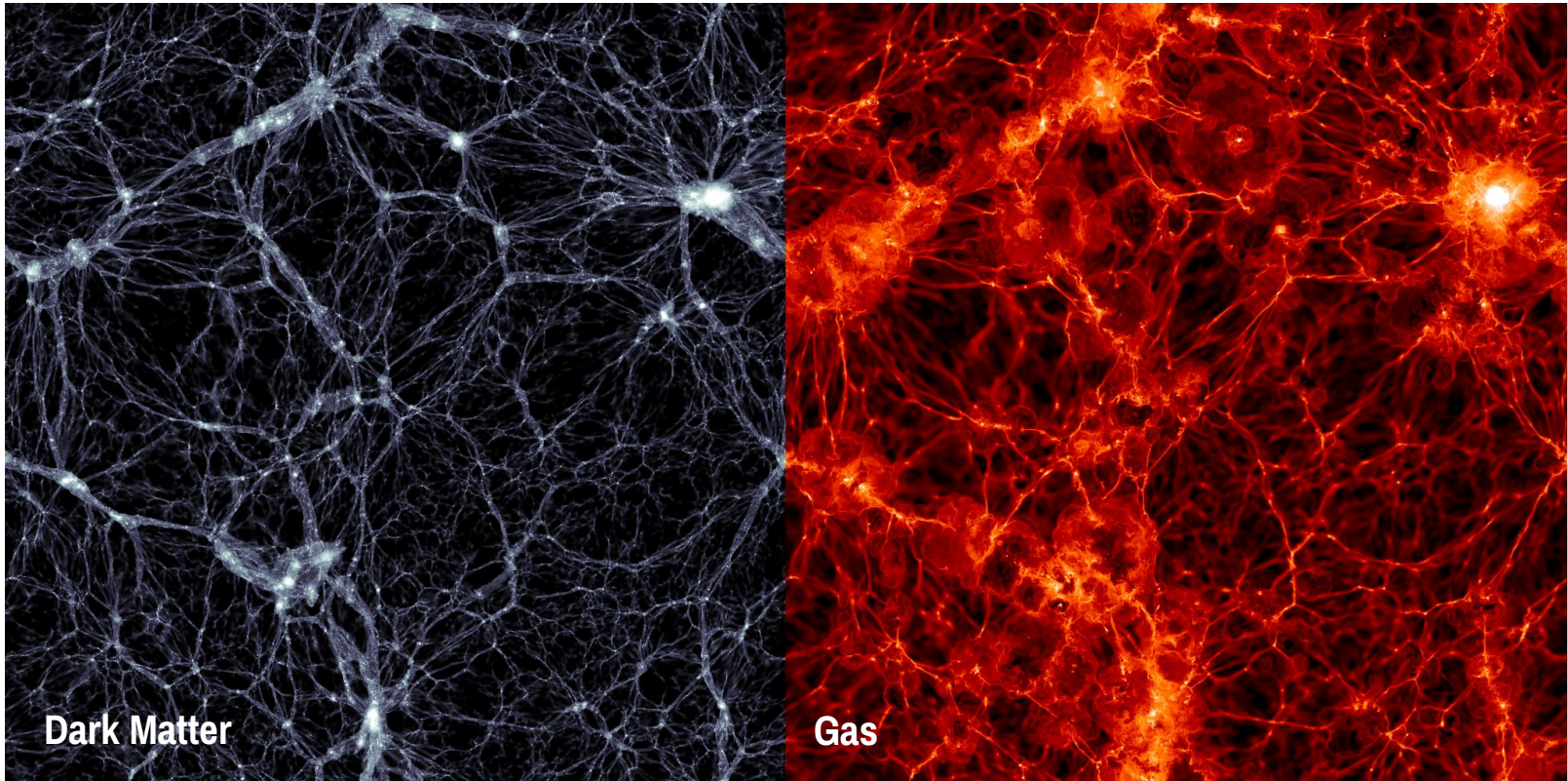


Adrian Bayer

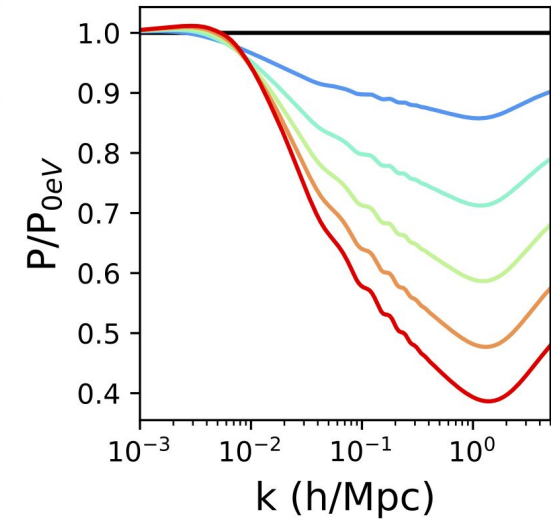
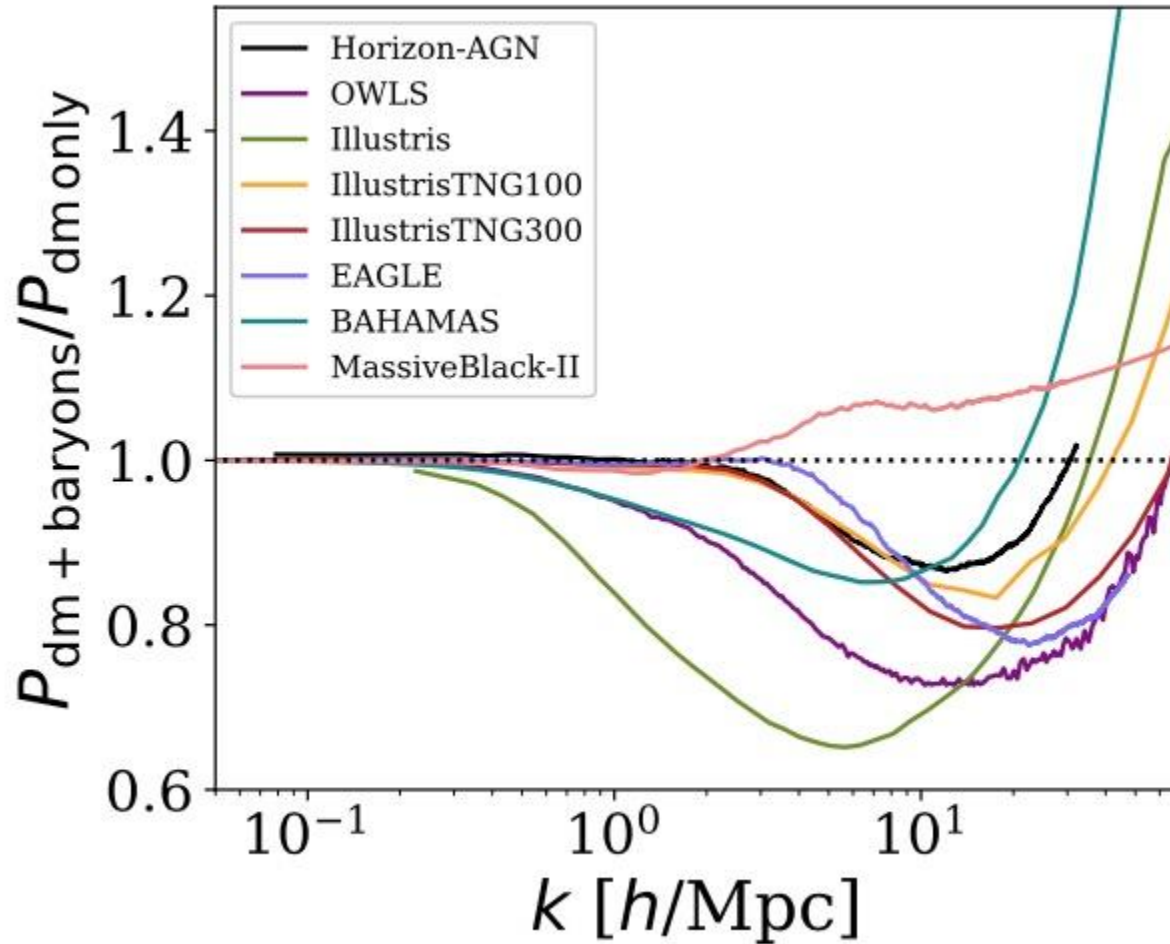
Bayer+(JL) 2021

Challenge 2: Baryons

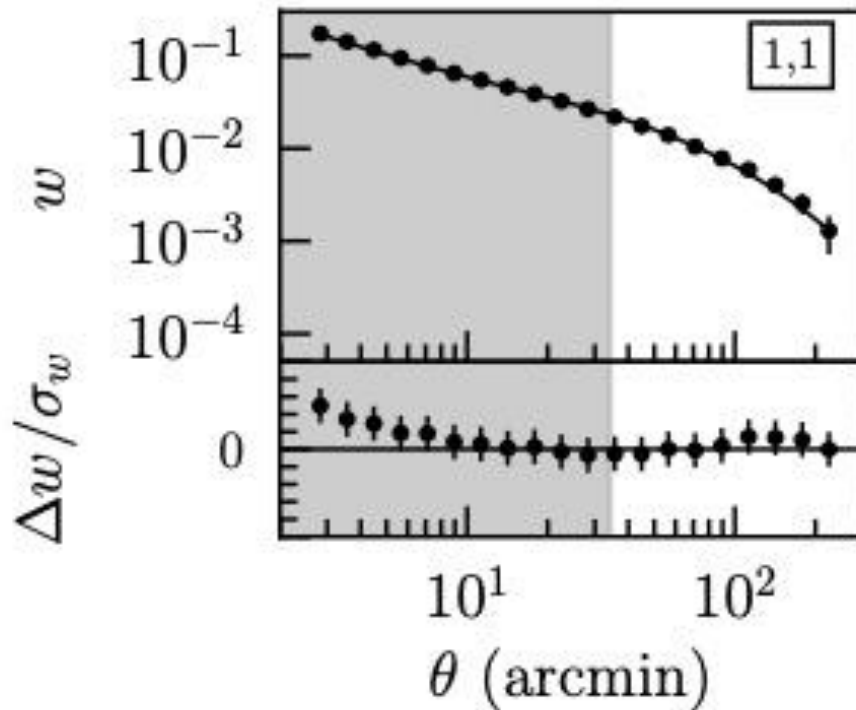
Baryonic Feedback



Like neutrinos, baryons also suppress the matter clustering



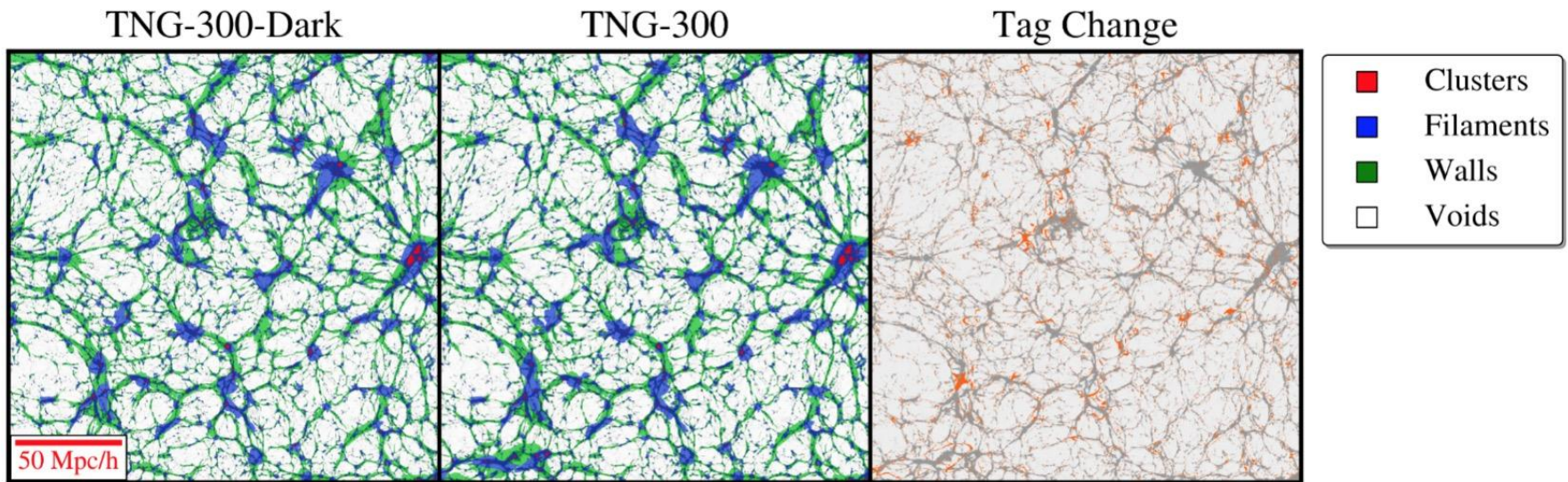
Already, we are limited by baryon models



The shaded regions indicate scales **not used** in the fiducial analysis, primarily due to uncertainties in the impact of **baryonic effects**.

Dark Energy Survey Y3 Results (2021)

*Angular correlation function of MagLim galaxies

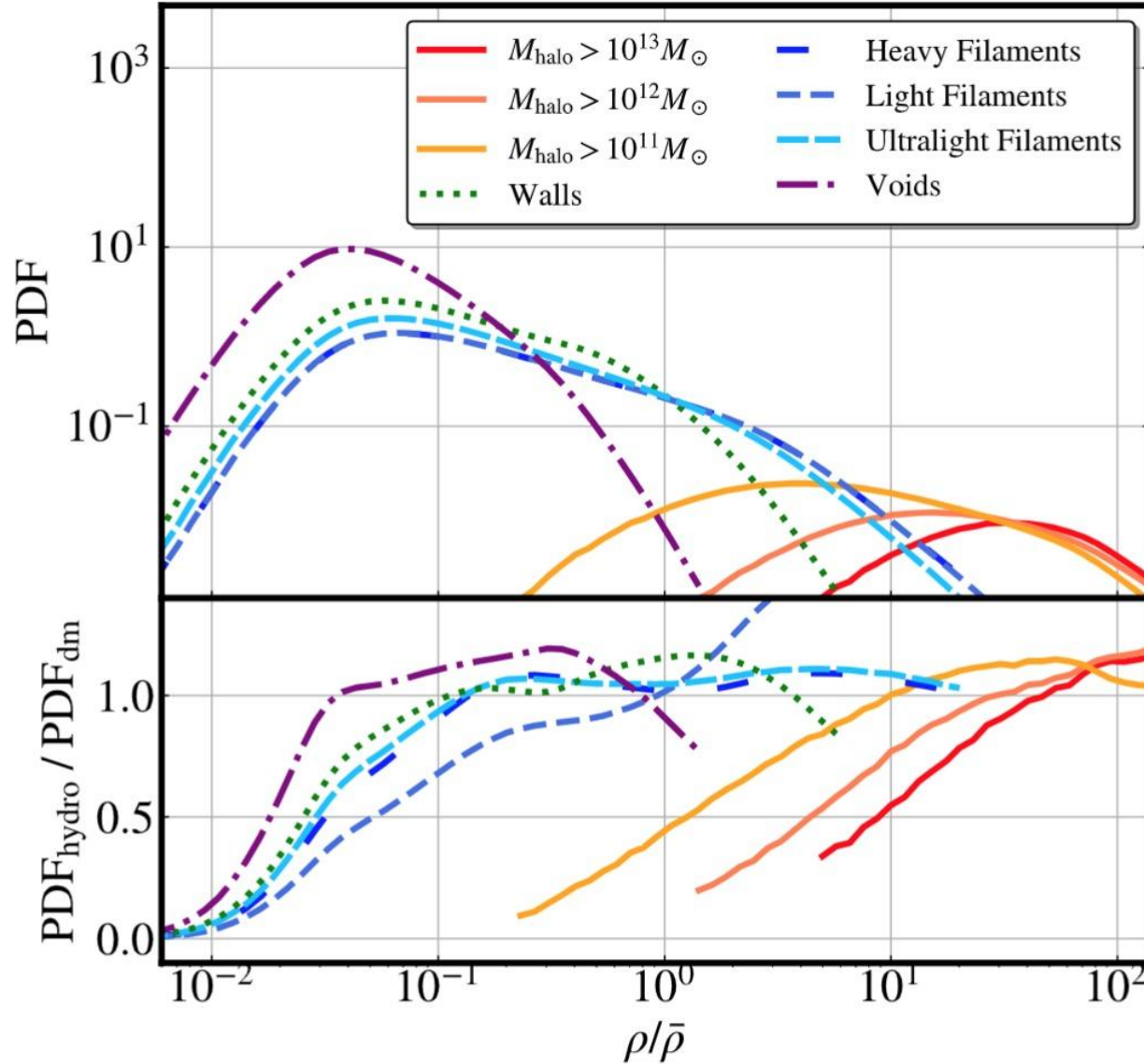


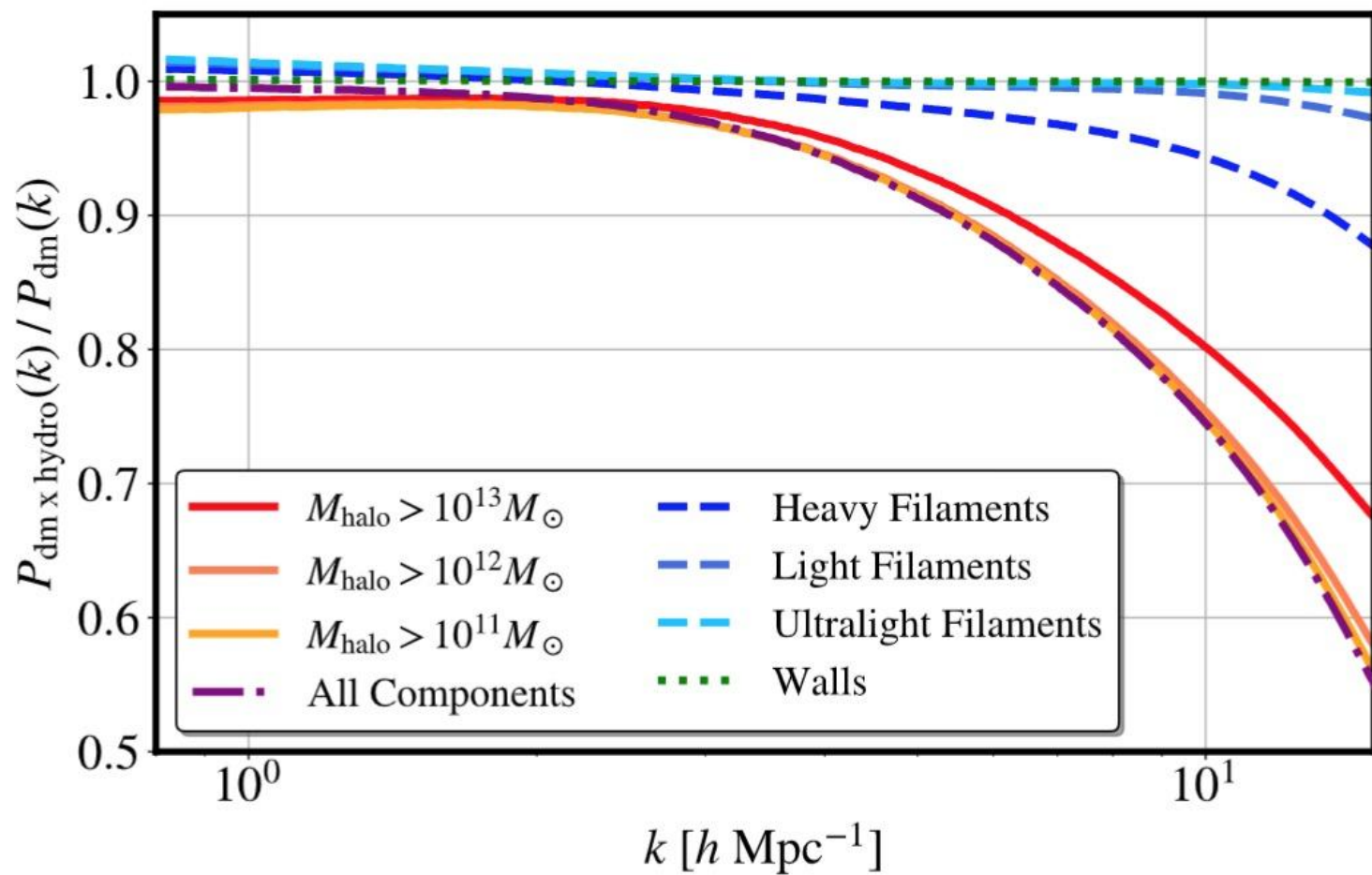
*Current baryon models focus mainly on halos. Here we examine the effect on **filaments, walls, and voids**.*



Sunseri, Li, JL (in prep)
Using the IllustrisTNG simulations

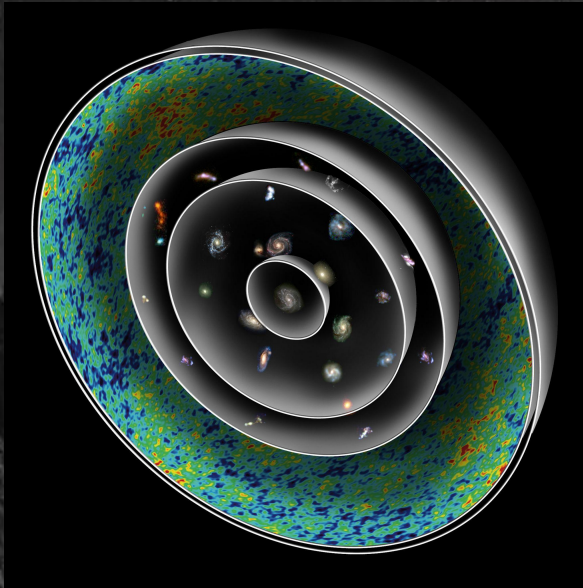
Baryons affects the cosmic web beyond halos





Challenge 3: Correlated CMB x LSS simulations

The Half Dome Simulation



Quick overview:

f_{sky} full
 M_{min} $10^{12} M_{\text{sun}}$
Box 5Gpc/h
 N_{particle} 8192^3
Release ~mid 2023

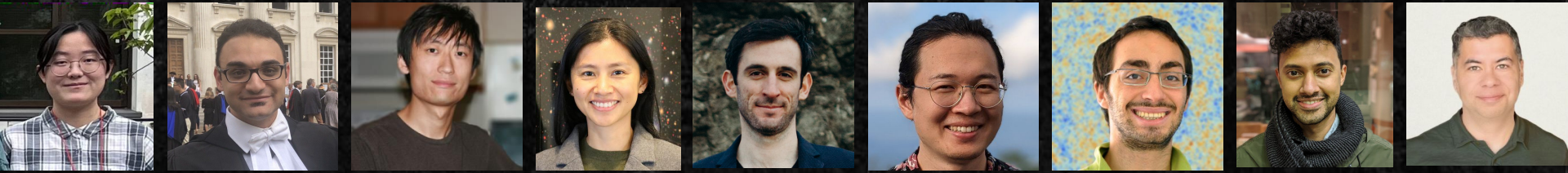
Healpix maps of:

tSZ, kSZ, CIB, radio,
CMB lensing, shear,
galaxies, clusters

N-body
linecones, particles,
velocities, halos, lensing

LSS
Shear, clusters, HOD...

CMB
tSZ, kSZ, CIB, Radio...



Yici Zhong Adrian Bayer Yu Feng Jia Liu Joe DeRose Zack Li Giuseppe Puglisi Mat Madhavacheril Marcelo Alvarez

Correlated Simulation Challenges

Assuming 1000 simulations*

1. Optical-CMB connections

AGN-Radio galaxies; SFR-CIB

Adding other observables (21cm, LIM, x-ray)?

2. Computing time: ~half billion CPU hours

Large memory need: quarter million cores simultaneously

3. Storage: ~10-100s PB

Raw: 500TB/run; aggressive downsampling needed->100TB per run

4. Cross-collaboration collaboration

Currently operate as a loose collaboration between people who are deeply embeded in DESI/LSST/SO/S4

Future access to multiple surveys' internal pipelines & proprietary data for future upgrades: N² MOUs??

Training and acknowledging simulation scientists

* A very conservative assumption..

Also see "Report from the Tri-Agency Cosmological Simulation Task Force" by Battaglia+2020

	Neutrino Mass	Mass Hierarchy
$\Sigma m_\nu > 0.1\text{eV}$	Discovery	No
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[SUMMARY]

(Almost guaranteed) discovery of neutrino mass or hierarchy in 10 years

New potential in the nonlinear regime, before LiteBIRD's tau (10 yrs)

Baryon models need to be improved beyond halos and 2pt

Correlated simulations face computational challenges (also opportunities!)