UCLA

Mergers and Collisions at the Heart of Galaxies

credit: JWST

Unsolved Problems in Astrophysics and Cosmology

A special thanks to Howard and Astrid Preston for their generous support



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Smadar Naoz

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The Densest Places in the Universe ~Every galaxy has a Supermassive Black Hole 10⁶⁻⁹M. Densest environments





Binary population Infer on other galactic nuclei Dense environment

The Nobel Prize in Physics 2020



III. Niklas Elmehed. © Nobel Media. Roger Penrose

III. Niklas Elmehed. © Nobe Media.

Reinhard Genzel



III. Niklas Elmehed. © Nobel Media Andrea Ghez

The central 10 arcsec ~ 0.4pc ~ 1.3 lightyear





Binaries near the monster in our backyard ◆ <u>3 confirmed Binaries</u>: Ott et al. 1999; Martins et al. 2006, Pfuhl et al. 2014. • Binary fraction of massive stars may be comparable to young clusters or larger (e.g., Ott et al. 1999; Rafelski et al. 2007; Pfuhl et al. 2014; Alexander et al 2008); Símílar to local OB eclipsing binary fraction Gautam et al.

• X-ray Binaries (e.g., Muno et al 2005a,b, Haggard et al 2017, Hailey et al 2018)

• Hypervelocity stars (e.g., Brown et al 2005,2008, Hills 1988; Miralda-Escude & Gould 2000; Yu and Tremaine 2003; Perets et al. 2009) • Stellar disk properties = Large binary fraction (Naoz et al 2018)





Densest Environment

Proxima Centauri

+ α Centauri A,B

Sun

~1.3pc, 4.2 lightyears





Densest Environment

Weak kicks with passing neighbors

Collisions with passing neighbors Collision



Timescale hierarchy



Sanaea Rose



For a stellar binary @ 0.5au



Rose, Naoz et al. 2020

Relaxation

Evaporation

Collision



 $\rho \sim r^{-\alpha}$

Sanaea Rose



For a stellar binary @ 0.5au



Rose, Naoz et al. 2020

Timescale hierarchy

Relaxation

Evaporation

Collision



Binaries at the heart of galaxies The Eccentric Kozai-Lidov (EKL) mechanism

inclination

Kozaí 1962, Lídov 1962, Naoz et al. 2011

Orbit normal

"inner"





Binaries at the heart of galaxies The Eccentric Kozai-Lidov (EKL) mechanism

circular

eccentric

20

t [yr]

10

50

0.1

10-3

10-4

 10^{-5}

GR effects: Naoz et al (2013)

[deg] 100 $m_1 = 10 M_{\odot}$ $m_2 = 1M_{\odot}$ $M_{SMBH} = 4 \times 10^6 M_{\odot}$ $a_1 = 10AU$ $a_2 = 0.003 pc$ _ 0.01 extreme $e_2 = 0.8$ eccentricity peaks cause mergers!

Naoz (2016) ARA&A arXiv:1601.07175

Compare to: "Standard" (quadrupole) Kozaí

30

40

Extreme eccentricities are common throughout the parameter space

Li, Naoz et al, (2014), ApJ 785, 116 + ApJ 791, 86

Teyssandier, Naoz, Lizarraga Rasio (2013), ApJ 779, 166



Binaries at the heart of galaxies + EKL (eccentric Kozai-Lidov) + General relativity (1PN) + GW Sgr A



+ Tides + Post main sequence stellar evolution (single and binary) + Unbinding the binary (fly-by) (see Rose, Naoz et al 2020) + Disruption due to the SMBH + Updated binary stellar evolution for solar and sub solar metallicities (e.g., Breivik et al 2019)

Hoang, **Naoz** et al (2018,2022) Stephan, Naoz et al (2016,2019) Wang, Stephan, Naoz et al (2021) Bao-Minh Hoang Huiyi (Cheryl) Wang Alexander Stephan





Binaries at the heart of galaxies Formation of G2-like objects





Witzel ... Naoz et al (2014) Gillessen et al (2012,2018)

G2 object

G2 survived!

Many ideas: Murray-Clay & Loeb 2012; Miralda-Escude' 2012; Morris, Meyer & Ghez 2012; Phifer et al. 2013; Guillochon et al. 2014; Pfuhl et al 2015; McCourt, & Madigan 2016; Madigan et al 2017; Peißker et al. 2021...



Binaries at the heart of galaxies





Binaries at the heart of galaxies Formation of G2-like objects



Sgr A

vatory Laser Team



Stephan, Naoz et al (2016,2019)



Ciurlo... Naoz et al. 2020



Binaries at the heart of galaxies centric Kozai-Lidov) White dwarf bin.

relativity (1PN) + GW

in sequ

ary)

+ Disruption du

15-150 in LISA

*LIGO/Virgo/KAGRA: + Unbinding the BH-BH rate ~ 17-45 Gpc⁻³ yr⁻¹ hole bin.

Blac

BH-BH Rate $\sim 1-20 \text{ Gpc}^{-3} \text{ yr}^{-1}$ 1-20 in LISA

Hoang, Naoz et al (2018,2022) Stephan, Naoz et al (2016,2019) Wang, Stephan, Naoz et al (2021)

Stellar merger

tion (single

ose, Naoz et al 207

ated binary stenar evolution for sol metallicities (e.g., Breiniet al NS-BH bin.

NS-NS Rate $\sim 0.2-1 \text{ Gpc}^{-3} \text{ yr}^{-1}$

NS-NS bin

NS-BH Rate $\sim 0.2-1 \text{ Gpc}^{-3} \text{ yr}^{-1}$

cataclysmic variables



Evolve to compact object

merge

Rose, Naoz et al (2020) Hoang, Naoz et al (2018,2022) Stephan, Naoz et al (2016,2019) Wang, Stephan, Naoz et al (2020)

Breakup

Tough life for a binary...







Dense Environment





Collisions in a Dense Environment BH - stellar collision ← ▲



For efficient accretion

Sanaea Rose



Rose, Naoz, Sari, Linial 2022





Collisions in a Dense Environment BH - stellar collision For efficient accretion

og(M/M ₀

IMBH Mass,



Sanaea Rose



Rose, Naoz, Sari, Linial 2022





Collisions in a Dense Environment BH - stellar collision For efficient accretion



Sanaea Rose



Rose, Naoz, Sari, Linial 2022

106 og(M/M ₀ 10⁵ 10^{4} IMBH Mass, 10³ 10² 10¹





Collisions in a Dense Environment BH - stellar collision For efficient accretion



Sanaea Rose



Rose, Naoz, Sari, Linial 2022



Collisions in a Dense Environment BH - stellar collision For efficient accretion



Sanaea Rose



GW mergers with the SMBH EMRIs and IMRIs

Rose, Naoz, Sari, Linial 2022

(⁰ M/M ⁰) 10³ 10² 10²





Collisions in a Dense Environment BH - stellar collision + Wind accretion

2.00

1.75

1.50

125

1.00

0.75

0.50

0.25

0.00

Density

Probability



Sanaea Rose Relaxation

 $\Gamma \sim 1.5 \text{ Gpc}^{-3} \text{ yr}^{-1}$ e.g. O'Leary et al (2009) Gondan et al (2018)

 $\rho \sim r^{-}$

Rose, Naoz, Sari, Linial 2022







Mergers and Collisions at the Heart of Galaxies

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Merging stars

Merging compact objects

Merging a compact object with a star





